

Error propagation in SMOS calibration subsystem

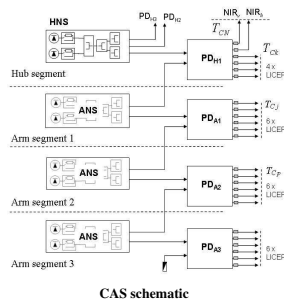
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- Ground characterization of Calibration Subsystem (CAS) of MIRAS (Microwave Imaging Radiometer using Aperture Synthesis)**
- CAS essential for overall performance of the payload
 - Ground calibration of CAS enables calibration of receiver gain, T_{SYS} and phase imbalance
 - The propagation and accumulation of errors in ground calibration to MIRAS calibration plane is simulated
 - An analytical method of calculating error propagation is introduced



Amplitude calibration
gain of hub segment receivers from LICEF readings v_k and ground calibration S-parameters of paths k, N

$$G_{PMSk} = \frac{v_{k,hot} - v_{k,warm}}{\left| \sum_{j=1}^N S_{kj} \right|^2 (T_{CN}^{hot} - T_{CN}^{warm})}$$

T_{CN} : noise levels measured by NIR at CAS outputs during noise injection.

→ gain of segment 1 receivers

$$G_{PMSj} = \frac{v_{j,hot} - v_{j,warm}}{\left| \sum_{k=1}^N S_{kj} \right|^2 G_{PMSk}}$$

Phase calibration
phase imbalance of receivers k and j solved with ground calibration

$$\theta_{kj} = \theta'_{kj} - (\theta_{ki} - \theta_{ji})$$

similarly, for receivers j and p

$$\theta_{jp} = \theta'_{jp} - (\theta_{ji} - \theta_{pq})$$

thus for any receivers k and p

$$\theta_{kp} = \theta'_{kp} - (\theta_{ki} - \theta_{ji}) + \theta'_{jp} - (\theta_{ji} - \theta_{pq})$$

Calibrated gain, T_{SYS} of receiver k ; 6 NIR readings used for T_{diff}

$$G_{PMSk} = \frac{v_{k2} - v_{k1}}{6 \sum_{i=1}^6 \left(\frac{v_{i,hot} - v_{i,warm}}{T_{diff}} \right)} = \frac{v_{k2} - v_{k1}}{6 \sum_{i=1}^6 X_{i,diff} T_{diff}}$$

$$X_{i,diff} = \frac{v_{i,hot} - v_{i,warm}}{v_{i,diff}}, \Delta X_{i,diff} = 2 \sqrt{\left(\frac{\Delta v_{i,hot}}{v_{i,diff}} \right)^2 + \left(\frac{\Delta v_{i,warm}}{v_{i,diff}} \right)^2} = 2\sqrt{2} \frac{\Delta v}{|v_{i,diff}|}$$

$$T_{SYS,k} = \frac{v_A}{G_{PMS,k}}$$

Uncertainty of G_{PMS} due to ΔX is:

$$\Delta G_{PMS,k} = \sqrt{6 \left(\frac{-T_{diff} T_{diff}^2 G_{PMS,k}}{\left(\sum_{i=1}^6 X_{i,diff} T_{diff} \right)^2} \Delta X_{i,diff} \right)^2}$$

assuming $X_{i,diff} = 1$, and $T_{diff,n} = T_{diff,k}$

$$\Delta G_{PMS,k} = \frac{1}{\sqrt{6}} G_{PMS,k} \Delta X_{i,diff}$$

→ Uncertainty of T_{SYS} Calibration

$$\Delta T_{SYS,k} = \frac{v_A}{G_{PMS,k}^2} \sum_{i=1}^6 \left(\frac{1}{G^2} \right) \Delta X_{i,diff}^2$$

finally

$$\Delta T_{SYS,k} = \frac{1}{\sqrt{6}} T_{SYS,k} \Delta X_{i,diff}$$

→ T_{SYS} expressed in terms of one amplitude measurement uncertainty

$$\frac{\Delta T_{SYS,k}}{T_{SYS,k}} = \frac{1}{\sqrt{6}} 2\sqrt{2} \frac{\Delta v}{|v_{i,diff}|} = \frac{2}{\sqrt{3}} \frac{\Delta v}{|v_{i,diff}|}$$

Phase imbalance of receivers k and p , in separate sections m and l , with section j in between

$$\theta_{kp} = \frac{1}{N} \sum_{i=1}^N (\theta'_{ki} - \theta'_{pi})$$

$$= \frac{1}{N} \sum_{i=1}^N (\theta'_k - \theta'_l - \theta'_l + \theta'_p)$$

$$= \frac{1}{N} \sum_{i=1}^N (\theta'_k + \theta'_p)$$

$$+ \sum_{i=1}^N \left(\frac{1}{N} \theta'_{k,i} - \frac{1}{N} \theta'_{l,i} + \frac{1}{N} \theta'_{l,i} - \frac{1}{N} \theta'_{p,i} \right)$$

θ_p and θ_k unaffected by measurement accuracy of CAS.

For N receivers, remaining terms sum up to $2N+2$ terms; assuming an identical error $\Delta\theta_{CAS}$ for values, RSS error in the calculated imbalance equals

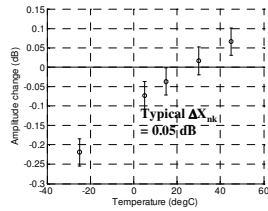
$$\Delta\theta_{kp} = \sqrt{2N \left(\frac{1}{N} \Delta\theta_{CAS} \right)^2 + 2(\Delta\theta_{CAS})^2} = \sqrt{\frac{2}{N} + 2} \cdot \Delta\theta_{CAS}$$

Calculation for multiple sections C between two receivers gives

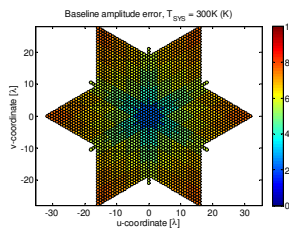
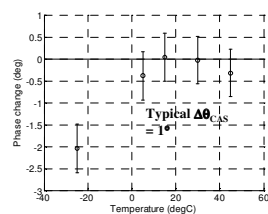
$$\Delta\theta_{kp} = \sqrt{\frac{2}{N} C + 2} \cdot \Delta\theta_{CAS}$$

CAS characterization over temperature

Amplitude change over temperature



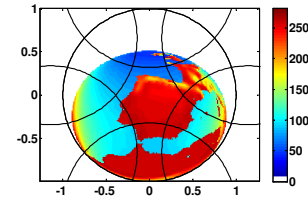
Phase change over temperature



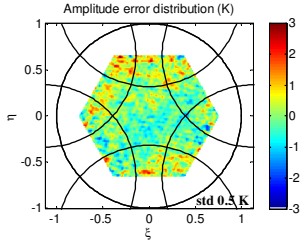
Calculated propagation of errors on MIRAS baseline chart

Standard deviation of error in alias-free zone

- 0.5 K for CAS amplitude errors
- 0.9 K for CAS phase errors
- combined effect in MIRAS image 1.0 K



Simulated MIRAS image using SEPS; including all error sources



Simulated MIRAS image using SEPS; effect of CAS amplitude errors